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Mechanics of Magma Chamber with the Implication of the Effect of CO₂ Fluxing

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Abstract

The rheological contrast between the viscous magma in the magmatic chambers and the surrounding rocks, having an elastic behavior, has allowed for a disturbance of the stress field, which, at a magmatic pressure unequal to the lithostatic one, can give rise to various modes of rock failure, leading to an eruption. In this context, one of the most important problems is represented by the mechanical stability of both large and extra-large magmatic chambers, such as the Yellowstone magmatic chamber. Due to its large volume, the critical overpressure necessary to start the volcanic eruption requires large added volumes of magma and fluids. The viscous relaxation of deviatoric stresses in the thermal areole of large chambers on a time scale of more than few years increases the critical volumetric flow rate of magma to 0.1–1 km³/yr. In this chapter, we have demonstrated that the deep CO₂ flux related with the underplating of basaltic magma has significantly enhanced the expansion of the magmatic chamber. In Yellowstone, the CO₂ fluxing can be the driver of its cyclic uplift and subsidence having a period of several decades. The extraction of water from the silicic melt by CO₂ increases the volume of the fluid up to 4 times, thus multiplying the pumping effect of the fluid. In a long term, simple estimates have also indicated that the CO₂ flux can significantly contribute to the heat balance, which can be up to the half one with respect to the value associated with the basaltic magma. The integrated stability index of a magma chamber has been tested during the beginning of the interglacial periods, when the rate of generation of the basaltic magma in a plume setting is of an order of magnitude higher than the normal one. The spikes of rhyolitic magmatism in some periods of the last Quaternary interglacials have been weakened for the Yellowstone case history. The last Pleistocene glaciation activated only strong hydrothermal eruptions, which may imply that at present, the level of CO₂ and basalt supply rate is not high enough to cause a major eruption in Yellowstone in the near future.

Keywords: magma chamber, viscoelasticity, CO₂, eruption, Yellowstone

1. Introduction

1.1. Basic relations

In a first approximation, a magmatic chamber can be regarded as the inclusion of a liquid in deformed elastic solid rocks. As a rule, viscous deviatoric stresses within this inclusion can be neglected and the stress state can be represented by hydrostatic pressure. The boundary between the enclosing rocks and magma can be approximately regarded as discrete. On this boundary, the continuity of the normal and zero tangential stress conditions are valid (e.g., [1]) or in vector notation:

$$\sigma \cdot n = P_m(t, x, y, z) \cdot n, \quad (1)$$

where n is normal to the chamber wall in point (x, y, z) , $P_m(t, x, y, z)$ is the magma pressure. The pressure of the magma depends on the deformation of the walls of the chamber, and also on the mass and state of the magma inside it [1]:

$$M(t) = \int_{V(P(t))} \rho(P_m(t, x, y, z)) dV \quad (2)$$

Differentiating the r.h.s. and l.h.s. of Eq.(2) with respect to time while using the space averaged pressure $P_m(t) = P_0 + dP_m(t)$ and expressing $V_{ch} = V_{ch,0} + dV(t)$, $M(t) = M_0 + q_m t$ and $\rho = \rho_0(1 + dP_m(t))$ yields another form of Eq. (2) which is used, for example, in Ref. [2]:

$$\frac{dV_{ch}}{V_{ch,0}} = -\frac{dP_m(t)}{K_m} + q_m dt / V_{ch,0} \rho_0 \quad (3)$$

where K_m is effective bulk modulus of the multi-phase magma, $V_{ch,0}$ and ρ_0 are reference magma chamber volume and magma density, q_m is mass flux of magma. When magma contains fluid bubbles its compressibility increases significantly [2]. Change of the magma chamber volume at the hosting rocks deformation can be expressed in effective form used in theoretical analysis as equation of state (dependence of the volume from the magma pressure P_m , which is a reference pressure in the apical part of chamber boundary when hydrostatic pressure gradient is included)

$$dV_{ch}/V_{ch,0} = dP_m(t)/K_{ch} \quad (4)$$

A chamber volume at the rising pressure increases not only due to the rocks compression as suggested, for example, in Ref. [2] but also due to the uplift of the surface and roof buckling. Therefore effective chamber bulk modulus K_{ch} can be at least order of magnitude smaller than that of the host rocks. Roof faulting reduces K_{ch} several times more [1]. In the correct numerical models boundary condition (1) and EOS of magma (2 or 3) are applied simultaneously (fully coupled system of equations) and pressure is determined iteratively at each time step.

The mechanical modeling of the magma chambers is mainly aimed at:

1. Prediction of localization and mode of the failure of the host rocks including chamber roof, volcano cone or in more general term magmatic—tectonic coupling;
2. Localization of the transport of magma out and into the chamber by dykes;
3. Simulation of processes during a volcanic eruption or on the way to an eruption, including the mechanical effects of filling the chamber with melts and fluids, the magma outflow through dykes, the magma crystallization and the degassing.

1.2. Tectonic-magmatic coupling

A strong rheological contrast between the magma and the rocks leads to the localization of the strains around the magmatic chamber, so that the rock deformation (and failure) is coupled with the magma emplacement on a global scale, especially in an extensional regime [3]. In a compressional tectonic setting, the magma chamber attracts thrusts, as shown in analog experiments and explained with an analytical solution for a circular hole in a compressed solid [4]. Simakin and Ghassemi [5] have studied numerically in 2D the effect of an elliptical magma chamber at shearing and showed that when interacting with the magmatic chamber, the strike-slip faults have experienced offset, similar to the step-over pattern.

In some situations, the influence of a pressurized magma chamber on a regional stress field can be represented by the solution of the problem of a pressurized circular hole in an elastic plate subjected to uniaxial stress at infinity [6]. The calculated directions of the maximum compressive main stress (s_1) are initially perpendicular to the chamber walls and, with an increasing distance from the chamber, rotate to the far-field stress direction. The dykes propagate through magma fracturing along the s_1 directions and perpendicular to the least compressive main stress (s_2 in 2D) [7]. Taking into account the influence of dykes on the stress field makes theoretical prediction of the dyke directions almost identical to those observed for the West Peak intrusion, Colorado [6].

1.3. Influence of shape

The mechanical properties of the magmatic chambers beneath the volcanoes depend on the shape of chamber, its relative depth (ratio depth to width), the gravitational load of the volcanic cone over the chamber. In general, magma chambers are horizontally elongated, as can be judged by the shape of solidified intrusions [8]. However, small shallow chambers under volcanoes often have an isometric shape close to a spherical. There are several theoretical studies in which the interaction of such spherical chambers with the surface is analyzed.

1.3.1. Spherical and cylindrical chambers

A closed-form analysis (improving classic Mogi [9] solution) of the stresses around a spherical inclusion (3D) in an elastic half-space was presented by McTigue [10]. The model predicts that the conic dykes (magma fracturing) can be expected to initiate on a ring of maximum $\sigma_{\phi\phi}$ on the sphere's surface. The ring is delineated by planes (lines) drawn tangent to the sphere

through a point on the symmetry axis at the surface. An accurate analytical solution for cylindrical chamber (circular in 2D) in the elastic half space was presented in Ref. [11]. Based on their solution for direction of the most compressive principle stress, authors developed conception of the capture zone for the ascending magma. They suggested that the magma ascending from the generation level in the zone of the differential stress above 1 MPa, caused by the presence of magma chamber, moves along s_1 directions toward it. Overpressurized spherical magma chamber redistributes s_1 directions in its vicinity normal to its surface and thus attracts ascending magma.

1.3.2. Elliptic chamber

The distribution of stresses around elliptical chambers near the surface has been numerically studied by many researchers (see in Ref. [12]). The main features of such solutions (in 2D) are shown in **Figures 1** and **2**. Both the overpressurized (**Figure 1a**) and underpressurized (**Figure 1b**) elliptic chambers when interacting with a free surface create an arch pattern of the maximum differential stress (s_1 - s_2). This parameter is twice larger than the maximum tensional principle deviatoric stress used in Ref. [11]). At the overpressure (**Figure 1a**), the preferred paths of dykes start at the edges of the chamber. Only a small part of the s_1

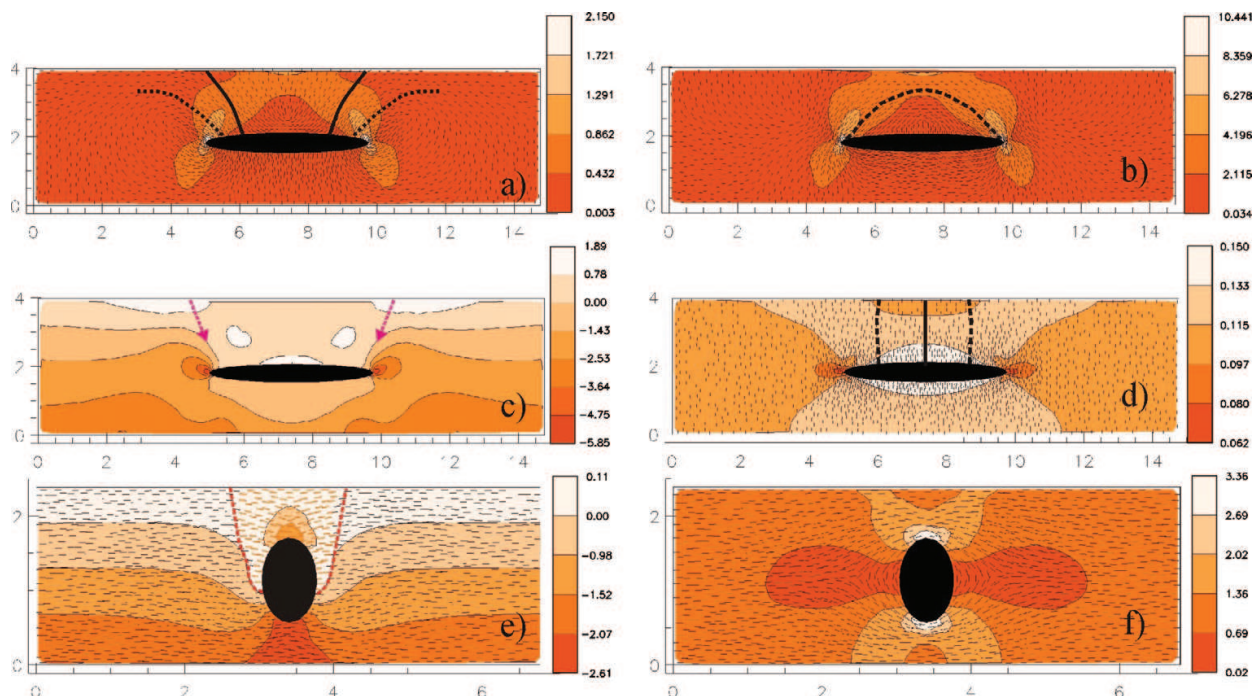


Figure 1. Examples of the calculated numerically in 2D elastic approximation of the stress fields around magma chambers without edifice. Spatial unit l_0 is arbitrarily chosen. Stress units are scaled with $\rho g l_0$, nondimensional Young modulus is 500 (at $l_0 = 2500$ m, $\rho = 2500$ kg/m³, $E = 30$ GPa), Poisson's ratio ν is 0.2. Density of the magma and rocks are equal. Short dashes map s_1 directions. (a) s_1 - s_2 , overpressurized chamber potential dykes trajectories are localized near the roof edges, (b) s_1 - s_2 in the underpressurized magma chamber, reference pressure in the apical point is zero, hydrostatic gradient exists, (c) s_2 in the underpressurized magma chamber, arrows depict normal faults starting on the surface, (d) s_1 - s_2 at extension, lines show potential trajectories of dykes out of magma chamber, (e) s_2 , overpressurized vertical chamber, potentially rocks within red dashed line can be ejected with diatreme formation, (f) s_1 - s_2 , underpressurized vertical chamber.

trajectories leads to the surface, while the other ones form the “saucer edges” known for shallow sills [13]. For the underpressurized chamber, the arch zone of the maximum shear stress (**Figure 1b**) can cause the detachment of the apical roof zone into the chamber. On the map of s_2 values (**Figure 1c**), one can note zones of the tensile stress on the surface above the edges of the chamber. Many researchers [12, 14, 15] suggest an essential role in normal inward dipping ring faults in the roof collapse and caldera formation. When the regional fault crossed the roof of the chamber in a position parallel to the arch of maximum s_1 - s_2 , the stress is amplified and this configuration can initiate the trap-door caldera formation at an enough overpressure [1]. During the extension, the conjugate normal faults tend to develop in the roof. The magma transport directions are vertical, and they all lead to the surface (see **Figure 1d**). A vertical chamber with excess pressure creates a stress field (**Figure 1e**), which can potentially lead to diatreme formation. The underpressurized vertical chamber only slightly disturbs the stress field (**Figure 1f**).

1.3.3. Elliptic chamber with cone

A perfect volcanic cone is often formed with the repeated effusive eruptions from the magma chamber. The cone loading significantly modifies the stress field. Without overpressure, the cone focuses the trajectories of the magma toward the apex of the cone (see **Figure 2a**). This theoretical prediction looks reasonable, since the huge perfect cones of Fujiyama and Klyuchevskoi volcanoes have a height of about 4 km above the base and were formed in hundreds of the repeated eruptions with magma outflow through the central crater at the top.

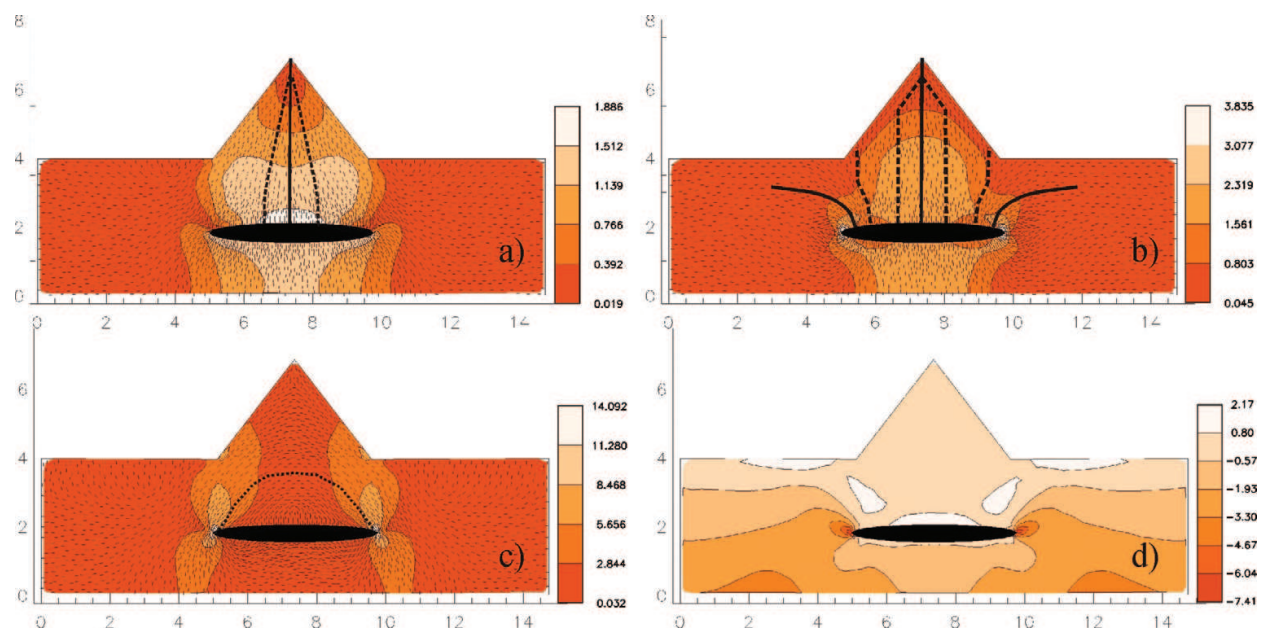


Figure 2. Calculated elastic stress field for magma chamber with edifice (calculations details as in Figure) (a) s_1 - s_2 uncompensated and unrelaxed respond on the large edifice load, magma is focused toward cone top, (b) s_1 - s_2 magma chamber is overpressurized, dykes trajectory toward the slope of cone appear, (c) s_1 - s_2 , reference $P_m = 0$, edifice slope instability is expected, (d) s_2 , reference $P_m = 0$, inward dipping normal faults formation is dictated by the extended linear zones with positive (tension) s_2 in the roof.

However, at a strong overpressure, focusing is weakened, and some trajectories can lead to the slope and to the base of the volcano (see **Figure 2b**). At the same time, Pinel and Jaupart [16] predicted that the dykes propagating vertically are pushed out from the edifice edges, probably because they did not include the magma chamber in their model. At the underpressure, the extensional zones develop on the surface and in the roof. Outward dipping zones of shear fractures also arise including the cone slope bottom (see **Figure 2c**). The failure of the bases of the cone can lead to the instability of the slopes of the volcanic edifice and to the consequent emplacement of debris avalanches, followed by an explosive eruption and by the formation of a horseshoe-shaped caldera. In this case, another failure mode is represented by extensional fissures dipping inwards (**Figure 2d**).

1.4. Effect of viscous relaxation of the deviatoric stresses

All the shown modes of the interactions of the magma chambers with the hosting rocks have been obtained in an elastic approach as an immediate reaction on the magma arrival or its evacuation. The long-term deformations around the magma chambers should be treated with accounting for the stress relaxation. The viscoelastic Maxwell rheology is appropriate for the modeling of the rocks with temperature above c.a. 400°C at enough long times. The characteristic time scale of a viscous relaxation is a ratio of the viscosity to Young modulus of the heated rocks $\tau_{ve} = \eta/E$, that is, 1–30 years at $\eta = 1\text{--}10 \times 10^{18}$ Pa s, $E = 10\text{--}30$ GPa. For example, geodetic data on the unrest in the Long Valley caldera in 1998 in the shorter time scale of several months were well interpreted accounting heated zones with viscosity $\eta = (1\text{--}100) \times 10^{16}$ Pa s [17].

In the two-dimensional numerical model by Simakin and Ghassemi [1], the viscosity of the rocks was calculated for the quasi-steady-state temperature field around an elliptic magma chamber. It was shown that at the constant influx rate (without magma loss into the dykes and at the eruptions), the magmatic overpressure (and corresponding deviatoric stresses in the hosting rocks) soon approaches an asymptotic value depending on the current magma chamber volume. This model is incomplete, since the paths leading to the probed current state of a chamber of a given size are not taken into account. Karlstrom et al. [18] have included the transient temperature field and the dyking in their complex model of magma chamber evolution. The mechanics was modeled with a rather complicated analytical solution for the circular inclusion in the viscoelastic half-space [18]. Obviously, many provisional features of viscoelastic shell were introduced due to the formal symmetry requirements. The model is not fully consistent, since the undefined initial chamber size has been used as a parameter in the evolutionary model. Melting and assimilation description is crude in comparison with special studies. Viscosity used in the calculations is not explicitly shown. Nevertheless by the order of magnitude results of both studies are quite close. According to Ref. [18], flux of about 0.004 km³/y is required for a chamber with volume 37 km³ to initiate eruption, similar parameters are obtained for the reference viscosity of the rocks in thermal aureole of $\eta = 5 \times 10^{19}$ Pa s at $T = 400^\circ\text{C}$ [1]. This value is within a range of the typical magma supply rates for continental silicic magmatic centers of $4.4 \pm 0.8 \times 10^{-3}$ km³/y. [19].

Both the models predict that for a large magma chamber, the magma fluxes as high as 0.1–1 km³/a are required to reach an overpressure capable to cause roof failures and caldera-forming

eruptions. In this chapter, we will consider additional factors that can influence the thermal and mechanical state of large magmatic chambers.

2. Deep fluid flux as an important mechanical factor

There is an increased interest in the giant silicic magma chambers, which in their life can become a batholith or a supervolcano, or both. Since in historical time, there were no supereruptions with a volume of 1000 km³ and more, judgments about the mechanisms of this phenomenon remain essentially theoretical. However, some aspects of this problem are almost completely clear. A huge volume of silicic magma can be formed by melting a silicic crust at the underplating with basaltic magma, which may include remelting of former rhyolitic volcanic rocks and granites. For the volcanic complex, Altiplano-Puna of the central Andes, including the caldera of La Pacana, with size 65 × 35 km, this mechanism is convincingly shown by geologic and geochemical data [20]. The Altiplano-Puna complex was formed at a thickening of the continental crust and intensive basalts generation in the mantle wedge at the steepening of the subduction angle and probably delamination of the crust [21]. For Bishop Tuff rhyolites (Long Valley, California), the composition of radiogenic isotope indicates a significant mantle contribution, which is interpreted as the consequence of the mid-crustal melting of amphibolites with basaltic protolith of mantle origin [22]. Another possible explanation for these observations may be the contamination of the silicic melts by REE, Pb and Sr, transferred by the CO₂-CO fluid from the underplating basalts. Below, we will consider some of the already established geochemical consequences of the fluxing of rhyolitic magma with deep CO₂. In any case, a sufficiently high rate of basalts generation is an essential condition for the production of a large volume of rhyolites. This requirement can be fulfilled in deep mantle plumes or in smaller upper mantle plumes (ascending flows) caused by the crustal delamination or breakage off the oceanic plate in the collision zones. The mechanics of assembling large volumes of silicic magma from small portions is analyzed in [23, 24]. In a long-term, magma in giant chambers is in a state of mush.

2.1. Why mush

2.1.1. Theoretical view

The melt in the partially molten zones of the Earth tends to separate from the solid phase when their densities are different. A two-phase melt-crystals homogeneous system is gravitationally unstable. At low concentration of the solid phase, the crystals separated from the magma by Stokes settling. When the solid-phase content becomes close to 40–45 vol.%, the crystals begin to interact mechanically and form a coherent matrix. At the melt density lower than that of matrix, melt separation takes place in the form of a compaction. The physics of compaction is quite complex and deserves special consideration, which is beyond the scope of this chapter. For the subjects considered below, it is important to know the stable content of the melt in the mushes of basic and silicic compositions. The effectiveness of the compaction is characterized by the compaction length δ_c [25]:

$$\delta_c = \left[\frac{\zeta + \frac{4}{3}\eta_s}{\eta_m} k(\varepsilon) \right]^{1/2} \quad (5)$$

where ζ and η_s are the bulk and shear viscosity of the matrix, respectively, η_m is magma viscosity, k is a permeability of the matrix with porosity ε . When the thickness of the mush layer is less than δ_c , the compaction is fast. In a detailed analysis of the matrix viscosity and permeability, Tegner et al. [26] estimated the compaction length for olivine-plagioclase cumulates of Skaergaard intrusion at an initial porosity of 22 vol.% at $\delta_c = 870$ m. Due to the fast compaction in the geologic sense with the mush thickness $< \delta_c$, the volume fraction of the residual melt in the olivine-plagioclase orthocumulates of the Skaergaard reduced to several volume percent [26]. This estimate is consistent with McKenzie's [25] conclusion that in the regions of basalt generation in plumes and in Mid-Ocean Ridges, the volume contents of the melt is no more than a few percent due to the rapid compaction and segregation of the melt into the veins and dykes. In silicic systems, compaction is less effective because of the low viscosity of a wet mush with a high quartz content and relatively high viscosity of the rhyolite melt. The corresponding meter-decimeter scale of δ_c and nonlinear rheological effects lead to local phase separation with the formation of a migmatite-like texture [27]. It is usually believed that additional tectonic deformations are involved in the separation of granitic magma and the formation of displaced granite plutons [28]. Thus, while in the basaltic system, partial melts with a volume fraction of melt greater than a few percent exist only temporarily [25], silicic mush, probably locally heterogeneous, with melt content of 10–20 vol.%, can persist for long time, limited by the rate of cooling and solidification.

The erupted Yellowstone intracaldera rhyolites are fairly homogeneous [23]. Periodic intrusion at different levels in the mush zone of superheated rhyolites from contact with underplating basalts can be a physical mechanism that wipes the mush heterogeneities and provides mixing on a large scale. The direct invasion of basaltic magma into the mush was hardly massive, since in Yellowstone only traces of clots of mafic minerals were identified in erupted intracaldera rhyolites [29], in contrast to the mixing textures well expressed in the extracaldera rhyolites [30].

2.1.2. Geophysical observations

A geophysical insight into the state of large silicic magma chambers requires a sufficiently large volume of observations. The seismic tomographic model of the Yellowstone volcanic system with two magmatic columns under resurgent domes merging at depth was proposed by Husen et al. [31]. It was significantly improved when data were combined from the dense seismic arrays of the Yellowstone, Teton and Snake River Plain (SRP) regional seismic networks, the NOISY array and the wide-aperture EarthScope Transportable Array ([32] and refs. in it). A continuous layer of a partial melt under the caldera at depths of about 6–12 km was clearly identified with a separate, presumably basaltic magma accumulation zone at depths 20–50 km. In this study, the volume fraction of the melt in the upper crustal body was estimated at 9%, and in the lower zone - 2 vol.%. Another independent source of information is provided by the magnetotelluric method. For the Yellowstone upper crust, the highly conductive zone of the hydrous partial melt is localized at approximately the same depths

6–12 km under the caldera [33, 34], as envisaged by seismic tomography. In fact, the interpretation of MT data is not absolutely unambiguous. High conductivity can be interpreted as a manifestation of a partial silicate melt or carbonatitic melt [35], aqueous or carbonic fluids and graphite. MT and seismic data may be complementary since the same volume fractions of carbonatitic and silicate melts will have a similar effect on the seismic velocity, while the conductivity of carbonatite melt will be at least two orders of magnitude higher [35]. Therefore, a very small volume fraction of carbonatites (or CO₂-CO-based solutions) will be detectable by MT and not visible in seismic data. The observed low resistance layer (seismically practically undistinguishable) under all SRP to the west of the Yellowstone caldera at depths of about 50–70 km can be a residual carbonatite melt with a low melting temperature or concentrated solution in carbonic fluid, rather than a basaltic partial melt.

In this chapter, we will focused on analyzing the large magma chambers created in some way such as the chamber under Yellowstone volcanic field which is one of the most studied (see e.g., review [36]). A generalized scheme of such a silicic chamber formed above the hot spot is shown in **Figure 3**. The possible role of volatiles in the heat budget and the mechanics of large silicic magma chambers is the least studied and will become the main subject of the following paragraphs.

2.2. High emission of deep CO₂

CO₂ is the main component that is released during degassing of basalts, especially from dry basalt formed during decompression melting in plumes. It was also assumed that in the lower crust there are reservoirs of the supercritical CO₂ derived in the mantle at 2–3 GPa [37]. A direct observation of the CO₂ mantle reservoir (at c.a. $P = 1.8$ GPa) sampled in the xenoliths from paleovolcanoes in the Pannonian basin is reported in Ref. [38]. In these xenoliths, CO₂/

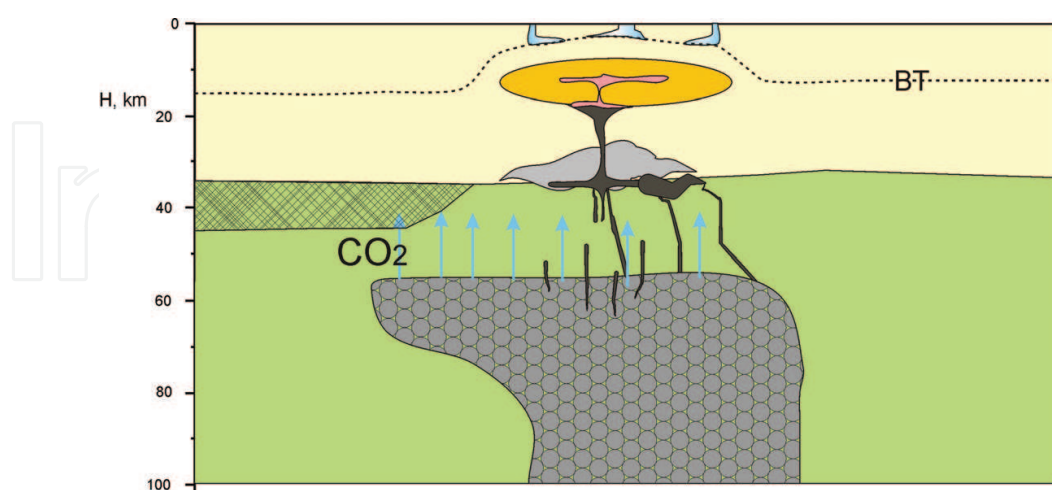


Figure 3. Cartoon schematically displays large silicic magma chamber in the continental hotspot setting. Basaltic magma originated in plume underplates crust (black) and where it partially solidifies to form mafic intrusions and partially rise to mush zone. At the lower contact of the mush zone, superheated rhyolite is formed and invades into mush. Hydrothermal system operates in the upper crust above brittle-ductile transition level. CO₂ infiltrates through the zone of magmatic activity and accumulates near the Moho away of this zone (crosshatched).

magmatic glass mass ratio is estimated of up to 0.25. The carbonated silicate melt with such CO_2 a content releases CO_2 at $P = 2\text{--}3$ GPa and later at their ascend, the basaltic magma and the CO_2 can move independently. Where deep sources of CO_2 exist, the magma chambers will be subjected to high fluxes of volatiles, which can play a significant role in keeping long-term heat balance [39], in modifying the composition of magma and in influencing the deformation of the magma chamber. Bachmann and Bergantz [40] attempted to model the essential water fluid Darcy's flow through the partial rhyolitic melt. In the model, water fluid was released from the basalt magma with 4–6 wt.% of dissolved H_2O . They found that possible fluid contribution in the instant mobilization of several 1000 km^3 of mushes immediately before eruption is negligible. Their conclusion is mainly determined by slow Darcy's filtration rate while faster mechanisms can exist.

In the following section, we will consider the effect of CO_2 fluxing on the example of Yellowstone caldera since the total CO_2 flux through Yellowstone caldera is among the largest and is about an order of magnitude greater than across the entire rift zone of Iceland [41]. Werner and Brantley [42] found that the largest flux of $456 \text{ kg/m}^2/\text{yr}$ in average is registered in the areas of the acid geotherms with surface 145 km^2 . Average CO_2 flux in the hydrothermal activity zones is about $40 \text{ kg/m}^2/\text{yr}$, while in inactive forestry area background flux is $3.5 \text{ kg/m}^2/\text{yr}$. Authors ascribe isotope composition of carbon and helium to 70% of CO_2 from the mantle source and the rest from sediments (Mammoth limestones). An analysis of CO_2 flux value showed that it is significantly larger than expected at the basalts magma degassing one at the rate of basalts generation in plume $0.03\text{--}0.05 \text{ km}^3/\text{yr}$ [42]. Extraordinary CO_2 generation can be linked with interaction of the mantle plume and Farallon plate. Part of subducted slab slides along $400\text{--}600 \text{ km}$ phase boundary under SRP [43]. It is separated by lateral tear from the older part of the slab intersecting this boundary. Plume crosses Farallon plate trough the tear. Presumably, oceanic slab contains carbonates that are mobilized by hot plume material and release CO_2 at pressure below $2\text{--}3$ GPa (see e.g., in Ref. [44]).

2.3. Heat budget with CO_2

The heat supply by deep fluids may be important in order to prevent from cooling the large volumes of the partial melts in the large and super-large magma chambers (such as the Yellowstone). In the most general form, the heat budget of the mush layer can be estimated adding the heat supplied by the different contributions of the carbonic fluid and basaltic magmas. The dissipation of heat depending on the specific physical mechanism can reduce the positive part of the balance. The heat lost is assumed to be essentially vertical, equal to the observed average heat flux, while the horizontal thermal loss is neglected. At the positive heat balance, the silicic melt will be generated and probably erupted, while at the negative balance the magma will solidify approaching a scenario of formation of batholiths.

The amount of the advective heat transfer by the ascending fluid depends on the mechanism of transport. If the fluid has passed the interval from contact with the basaltic magma to the layer of the partial rhyolitic melt along the fractures rather than by the Darcy's flow, then the heat losses through the crack walls will be small, and we can assume that the heat flux is approximately equal to

$$Q_{adv} = c_{p,fl}(1200 - 850)q_{CO_2} = 0.22e08 \text{ J/m}^2/\text{yr} \text{ (0.70 W/m}^2\text{)}, \quad (6)$$

where q_{CO_2} is mass flux of CO₂ in kg/m²/yr, the specific heat of the fluid $c_{p,fl}$ is taken equal to 1.4 KJ/kg/o, the temperatures of basaltic magma and rhyolitic mush are 1200 and 850°C, respectively. As the first proxy, $q_{CO_2} = 40 \text{ kg/m}^2/\text{yr}$ is taken equal to the average value of flux observed in the geothermal regions of Yellowstone [42]. Below we consider the effect of the variation in q_{CO_2} over a wide range on the heat balance.

In fact, the carbonic fluid exsolving from the underplated basalts and stored in the mantle reservoirs is a reduced. At $P = 2\text{--}3 \text{ kbar}$, the mole fraction of CO in such a fluid at f_{O_2} around QFM-0.5 is in the range 0.1–0.2 depending on the temperature [45]. At the oxidation of CO heat is released:



The enthalpy ΔH_T of the oxidation reaction (7) at $T = 700^\circ\text{C}$ is -198870.4 J/mol . Volume effect of the reaction (7) is negative, so the heat effect is somewhat larger than ΔH_T at a pressure $P = 2 \text{ Kbar}$ corresponding to a chamber depth of 7 km. However, not pure oxygen can be involved in the reaction, so we use the value of ΔH_T as the maximum estimate. If we take q_{CO_2} equal to $40 \text{ kg/m}^2/\text{yr}$, as above, and set $X_{CO} = 0.1$, then the heat of CO oxidation will be equal to

$$Q_{CO} \approx -0.1 \Delta H q_{CO_2} 1000/44 = 0.18e08 \text{ J/m}^2/\text{yr} \text{ (0.57 W/m}^2\text{)} \quad (8)$$

The main source of heat in Yellowstone is associated with the flow of basaltic magma. The simple arrival of $q_b = 0.05 \text{ km}^3$ of magma annually means advective heat flux

$$Q_b = q_b(\rho_b c_{p,b}(1200 - 850) + \Delta H_{m,b})/S_0 = 0.51e08 \text{ J/m}^2/\text{yr} \text{ (1.62 W/m}^2\text{)} \quad (9)$$

where the surface of the active part of the caldera is roughly estimated at 2000 km^2 , the density of basaltic magma is set 2700 kg/m^3 , the specific heat $c_{p,b}$ is 1 kJ/kg/o and the latent heat of fusion, absorbed at melting of rhyolites is $\Delta H_{m,b} = 400 \text{ kJ/kg}$. With these values of the parameters, the value of the advective heat flux Q_b is comparable in order of magnitude with the heat flux associated with CO₂.

Basalts generally do not intersect partial rhyolite layer serving a density filter. Their heat content is transferred to the superheated rhyolites formed on the lower boundary of mush zone that can penetrate in the partial melt. While fluid is not accumulated in the crust and rise through partial melt transferring heat and dissolved components into and out of the mush zone. Purely conductive heat transfer through partial melt and upper thermal aureole of magma chamber is limited by the stationary heat flux through mush layer in the absence of convection of the order

$$Q_{cond} = \lambda \Delta T / H_{mush} \quad (10)$$

where $\lambda = 2.5 \text{ W/m/o}$, $H_{mush} = 6 \text{ km}$. With these parameters values, the conductive heat flux will be only 0.145 W/m^2 or $0.05e08 \text{ J/m}^2/\text{yr}$. The average heat flux through Yellowstone caldera is significantly larger than this conductive estimate.

The dense basaltic magma interacting with the lower contact of the mush zone transfers its heat content to the superheated rhyolites formed during this interaction. The superheated rhyolite penetrates into the mush zone and indirectly transfers Q_b . The low density fluid moves directly through the lower contact into a partial melt and transfers heat and dissolved components.

2.4. Thermal profile through the mush zone

The above global estimates are valid regardless of the real physical mechanism of fluid transport. A self-consistent mechanical model of fluid transport through a partial melt should include the coupled magma and fluid flow through a deformable viscoelastic matrix that extends the viscous [25] and visco-pastic [27] compaction models. With a sufficiently comprehensive model one can consider complex flow regimes similar to a self-organized critical transport observed experimentally [46]. Below, we consider the simplest estimate of the spatial temperature distribution in a mush with a pervasive fluid flow as a solution of the general advection problem in a porous media. The mass flux of the fluid will not be found as a part of the solution (e.g., as a solution of Darcy's equation, as in Ref. [40]) but is specified on the basis of observations. In the general form, 1-D equation of advective heat transfer through a porous medium is

$$\rho_s c_{ps} \frac{\partial}{\partial t} T(z, t) = \lambda \frac{\partial^2 T(z, t)}{\partial z^2} + c_{pf} Q_f \frac{\partial}{\partial z} T(z, t) \quad (11)$$

or in nondimensional form

$$\frac{\partial}{\partial t} \bar{T}(z, t) = \frac{\partial^2 \bar{T}(z, t)}{\partial z^2} + Pe \frac{\partial}{\partial z} \bar{T}(z, t) \quad (12)$$

where Peclet number is $Pe = l_0 c_{pf} Q_f / k_T \rho_s c_{ps}$. Substituting $l_0 = 200$ m, $k_T = 1.0e-06$ m²/s, $c_{pf} = 1.4$ kJ/kg/o, $c_{ps} = 0.84$ kJ/kg/o, $\rho_s = 2500$ kg/m³, we get the time scale 1300 yrs, $Pe = 0.013-0.4$ with q_f in the range 4-100 kg/m²/yr. At this time and spatial scales problem is weakly to moderately advective ($Pe < 1$). It is important to set proper boundary conditions for temperature.

At the lower boundary, the temperature of basaltic magma $T_b = 1200^\circ\text{C}$ is prescribed. The heat flux from the mush layer supports the heat flux in the geothermal system and in the quasi-steady state their values should be close in average by area and time (as used in 1D models). In our calculations, a heat flux Q_h of 4.6 GW/2000km² = 2.3 W/m² at the upper boundary is set, equal to the average estimates for Yellowstone [36].

We calculated the transient solutions for different Pe values. At $Pe = 0.13$, corresponding to a mass flux of CO₂ of 40 kg/m²/yr in a geologically short time of 13 Kyr, the temperature at the top of the mush layer drops to 200°C and continues to decrease and the thickness of the mush layer is halved (see **Figure 4a**). With a stronger fluid flow $Pe = 0.5$ in time equal to 56 Kyr, the boundary temperature still increased to c.a. 500°C, but did not reach a steady-state value (see **Figure 4b**).

In our formulation, in steady state, there is relationship between $Pe(q_{fl})$ and the maximum admissible heat flux $Q_{h,max}$. This is determined by the fact that temperature at the outer

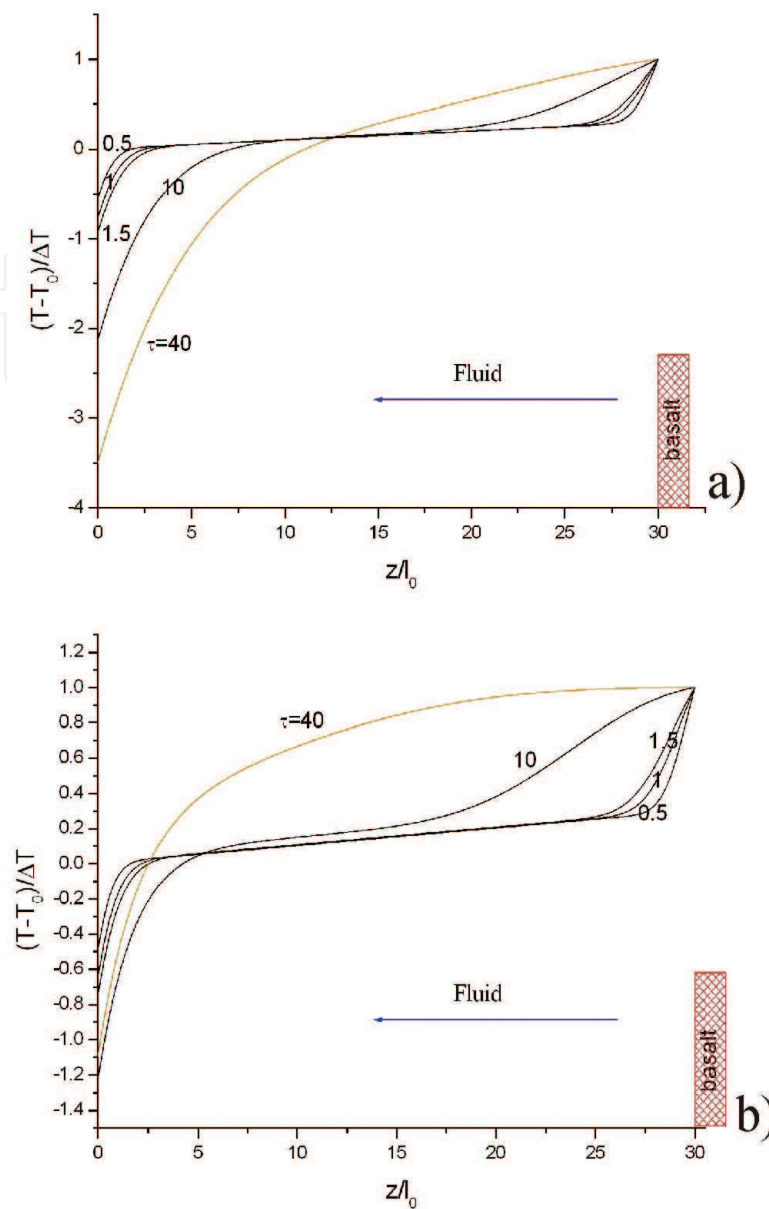


Figure 4. Calculated temperature profiles in the mush zone in nondimensional form in the moments of time $\tau = 0.5, 1.0, 1.5, 10$ and 40 . Length scale $l_0 = 200$ m, time scale $t_0 = 1300$ years. Right boundary is at $T = T_b = 1200^\circ\text{C}$, left boundary is subjected to the constant heat flux of 2.3 W, initial temperature of mush $T_0 = 850^\circ\text{C}$, (a) $Pe = 0.13$ (corresponds to $q_{\text{CO}_2} = 40$ kg/m²/yr, (b) $Pe = 0.5$. In (a) at $\tau = 40$ temperature at left boundary is physically meaningless, since constant heat flux boundary condition does not guarantee even positive absolute temperature. At $Pe = 0.5$, left boundary temperature stabilizes around reasonable $T = 500^\circ\text{C}$, however then almost one third of the mush zone asymptotically is heated to T_b .

boundary (closer to surface) is part of solution. We fix a reasonable value of $T(0) = -0.7$ (about 600°C in physical dimension) and get

$$Q_h = 1.7Pe / (1 - \exp(-30.0Pe)) \quad (13)$$

The numerical solution agrees with Eq. (13), at $Pe = 0.13$ the maximum $Q_{h,\text{max}} = 0.226$ (Eq.13) is less than 0.73 , used in the numerical solution, and at $Pe = 0.5$ $Q_{h,\text{max}} = 0.85 > 0.73$. As expected the mush layer solidifies in the first and melts in the second case, respectively.

The heat flux is measured directly only on a small part of caldera. The measurements of the heat flow at the bottom of Yellowstone Lake [47] gave values 100–300 mW/m² near the caldera boundary and up to 1600 mW/m² in Mary Bay. Fournier [48] obtained an estimate of 2.65 W/m² (with an active surface 2000 km²) on the basis of the correlation of the chlorine content and the enthalpy of the geothermal fluid. A later inventory of the chlorine balance in the Yellowstone river system [49] gives an average heat flux of 2.3–3.3 W/m² (when recalculated to an active caldera area $S = 2000 \text{ km}^2$). The CO₂ flux was measured in geothermal areas at a quiet stage of the caldera evolution. The amount of CO₂ accumulated and periodically released is unknown. In fact, deep CO₂ is partially redistributed horizontally, accumulated underground and transferred as an HCO₃[−] ion by geothermal waters. The rate of basaltic magma supply (q_b) is also known with low accuracy. In view of all these uncertainties, we are considering the heat budget in a wide range of variation of the parameters involved and display the results in **Figure 5**.

The zero heat balance conditions in the mush zone for different q_b are indicated by solid lines in the coordinates $q_{\text{CO}_2} - Q_h$ (surface heat flux). In general, the system is close to the global balance at $q_b = 0.05 \text{ km}^3/\text{yr}$. Doubling of this rate shifts it to the rhyolite magma generation. Variations in the flux of CO₂ can lead to the transitions from solidification to melting in a state close to thermal equilibrium. The dashed line shows the dependence of the heat and CO₂ fluxes calculated using Eq. (13). Evidently, when basalt is passively underplating the mush zone, providing $T = T_b$ (boundary condition in the advection model) only an extremely high CO₂ flux can ensure the long-term existence of the mush layer. This analysis emphasizes the importance of the interaction of mush with superheated dry rhyolites (see **Figure 3**). In fact,

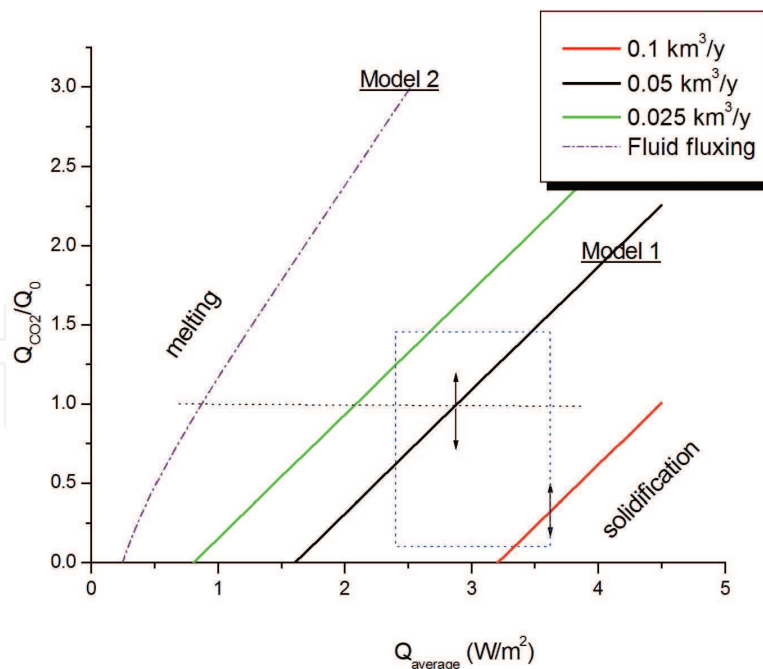


Figure 5. Zero heat balance condition for rhyolitic mush zone of Yellowstone volcano in the coordinates average surface heat and CO₂ fluxes (a) continuous lines correspond to the different basalt magma supply rates and mechanisms providing full heat transfer of basalt magma and deep CO₂ fluid to the mush, (b) dash-dot line plots relationship for advective fluid heat transfer through the mush and passive basalt underplating setting low boundary constant temperature $T = T_b$ condition. CO₂ flux is normalized on the $Q_0 = 40 \text{ kg/m}^2/\text{yr}$. At the heat flux larger than zero-balance value mush is solidified at the smaller heat flux or larger CO₂ flux mush is melted.

convective melting and mixing is fairly fast process [50], so that zircon crystals from melted mush will survive. Mineral thermometers will record crystallization temperatures at a solids content of about 40 vol.%, when convective melting stops, well below the initial temperature of the superheated rhyolites.

2.5. Volatiles in the melt at CO₂ fluxing

Our analysis demonstrates that the flux of CO₂ in a reasonable range can make a significant contribution to the heat budget of the mush layer. When crossing a layer of mush, the fluid should create petrologic signs of interaction with hydrous rhyolitic magma, for example, by transferring water from the melt to the CO₂ bubbles. During ascent and eruption, volatiles escape from the bulk melt, information about the pre and sin-eruptive conditions is retained in the melt inclusions in magmatic minerals. In the equilibrium with the fluid, the contents of H₂O and CO₂ in the melt depend on the PT conditions and the composition of the fluid. We approximate the mutual solubility of CO₂ and H₂O in the rhyolitic melt, using the compilation of experimental data from Ref. [51]. The solubilities in the melt are expressed through the pressure (in Kbar) and the mole fraction of CO₂ (*z*) in the fluid as:

$$C_{H_2O} = 4.1\sqrt{P(1-z)^{1.35}}, \quad C_{CO_2} = 0.0499kPz^{0.75}, \quad k = -0.54013 + 0.0164T, \quad (14)$$

where *C* is in wt.%, temperature in °C. The widely used model of the mutual CO₂-H₂O solubility [52] in the rhyolitic melt is more accurate, but also more complicated and can not be incorporated in numerical codes, such as a two-phase magma flow simulator. Our short expressions are for these numerical applications.

To model the melt and fluid equilibria, we solve the equations of mass balances for CO₂ and H₂O along with Eqs. (15). In the **Figure 6a**, the isobar for *P* = 2 Kbar for the composition of the rhyolite melt in coordinates *C*_{H₂O}–*C*_{CO₂} is plotted, the maximum concentration of water and CO₂ are for the pure water and CO₂ fluid, respectively. Various processes affecting volatiles in magma have been modeled and are depicted in **Figure 6a**. When the melt is depressurized during the eruption, degassing begins with the shift of the melt composition along the trajectories of degassing under the open (when all exsolved fluid is removed) or close system. Trajectories are shaped by the early loss of CO₂, followed by the release of H₂O (more on this in Ref. [52]). Melting of the mush as a close system leads to dilution of the melt with respect volatiles (movement of the composition from the initial point to the origin of coordinates in **Figure 6a**). Crystallization at constant pressure leads to a shift to more water-rich compositions along the isobar. If the fluid exsolved during crystallization remains in the mush, the pressure rises. The result of mixing with superheated rhyolite depends on the composition of the later. A transition to water poor, fluid undersaturated compositions is expected in general. At CO₂ fluxing at constant pressure, the composition follows isobar from the water-rich to water-poor melt. During CO₂ flushing, the pressure of the magma can increase due to the expansion of the chamber being constrained by the host rocks. Then the trajectory of the melt composition in H₂O-CO₂ space will deviate from the isobar (see **Figure 6a**). A significant vertical size of the magma chamber can also cause deviations of the MIs compositions from the isobar [53].

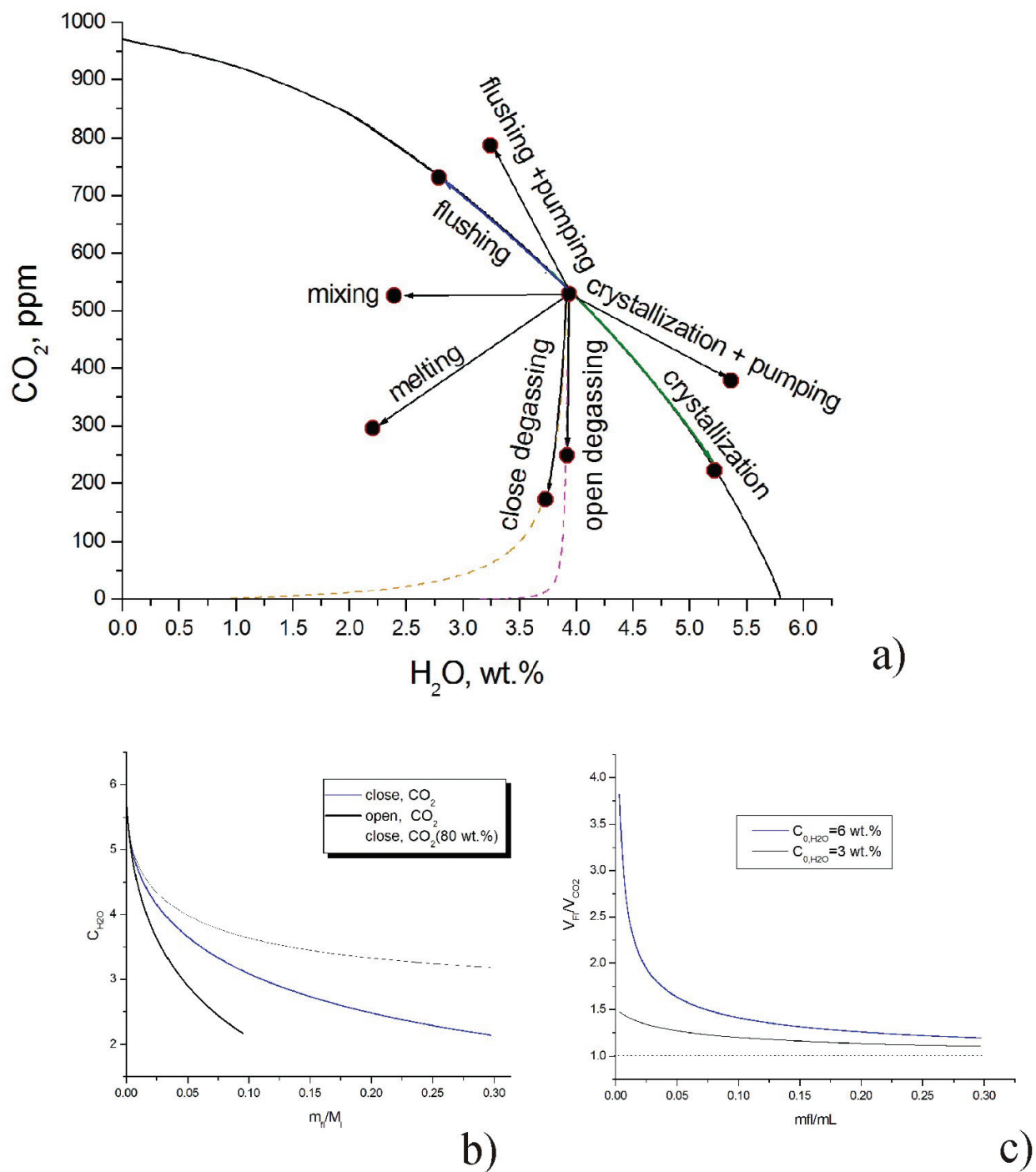


Figure 6. Results of modeling (a) different factors affecting volatiles content in the melt, (b) dependence of the water content in the magma at the flushing with CO_2 on the mass ratio of the reacted CO_2 and melt. Style of the interaction and content of CO_2 in the incoming fluid are shown in the legend (c) ratio of the volume of CO_2 - H_2O composition fluid to the initial volume of CO_2 as function of the mass ratio of the reacted CO_2 and melt. In the legend initial water contents in the melt are shown.

The parameters of the CO_2 flushing process require more detailed consideration, since they are necessary for further mechanical analysis. The modeling of the closed system flushing was carried out by increasing of the total contents of CO_2 and H_2O in the system in small

increments in the proportion of the fluid composition. To evaluate the volume of a fluid in an equilibrium two-phase system, we use EOS for a CO₂-H₂O mixture from Ref. [54]. In **Figure 6b**, one can see that added mass of the fluid increases, the concentration of water in the melt gradually drops. Pure CO₂ eventually will extract all the water from the melt. With a starting content of 6 wt.%, it is sufficient to add about 30% of CO₂ by mass to reduce the water content to 2 wt.%. In the open system approximation, after every step, all the fluid is removed, and a new portion of carbonic fluid reacts with the pure melt. In this case, approximately 10 wt.% of CO₂ is sufficient to achieve the same effect. When the incoming fluid contains 20 wt.% of H₂O, it is equilibrated with the melt with c.a. 3 wt.% of H₂O and cannot reduce the water content below this value. This implies that in the SRP system, carbonic fluid penetrating the rhyolite mush contains less than 20 wt.% of water to produce melts with a lower content. The transfer of water from the melt to the fluid changes the composition of the fluid and enlarges its volume. In **Figure 6c**, the volume ratio of the final fluid and the incoming pure CO₂ at P = 2 kbar is plotted with the two initial water contents. A strong volume increase of 1.5–4 times has place at the lowest mass ratio of the added fluid to the melt. As the mass ratio increases, the volume expansion becomes less significant. Extraction of water and CO₂ dissolution in the melt at the constant temperature induces crystallization due to an increase in the liquidus temperature. Crystallization caused by water loss will release heat of fusion and heat the melt, so that the integral effect should be more precisely evaluated with numerical modeling, which is beyond the scope of this chapter.

2.5.1. Melt inclusions data supporting the concept of CO₂ flushed mush

Figure 7 shows the results of the studies of melt inclusions in quartz phenocrysts from various large-volume rhyolites from Snake River Plain (SRP) and Long Valley calderas (Western USA). Lava Creek Tuff (LCT, Yellowstone 0.64 Ma) [55] decompression sequence is rooted in a 2 kbar isobar with an initial relatively high water content of 3 wt.%. The post caldera Tuffs of Bluff Point (Yellowstone, TBP 175–180 kyr) [56] sequence corresponds to degassing at the decompression from a pressure slightly higher than 1.5 kbar with a starting water content of 2 wt.%. Early Arbon Valley Tuff (AVT) [57] represents the products of the Picabo caldera eruption (10.44 Ma) with the highest water content of 6.5 wt.%. The AVT-LCT-TBP sequence illustrates a gradual decrease in water and an increase in the CO₂ content. Bishop Tuff (BT) MIs [53] demonstrate a highly variable set starting in CO₂ free compositions with c.a. 6 wt.% of H₂O. Some points follow approximately an isobar of 2 kbar, while others lie on the open system degassing path. One exceptional point with the highest CO₂ content reaches an equilibration pressure of more than 2.5 kbar. In general, the BT data are consistent with the fact that the magma was flushed with CO₂ followed by an eruption. Alternatively, the observed CO₂ enrichment and H₂O loss process can be explained by mixing superheated rhyolites with a high CO₂ content and hydrous partial melts (see **Figure 6a**). It is likely that AVT data with a simple shift of compositional points along the H₂O axis may correspond to such a mixing process that creates a fluid undersaturated melt. For comparison, we plotted in **Figure 7**, the MIs in quartz data for the Toba volcano [58] located in the subduction zone. Obviously, in a subduction setting with much lower CO₂ fluxes and a shorter residence time, the magma contains much less CO₂. However, even in this case, one can expect a CO₂ fluxing with the

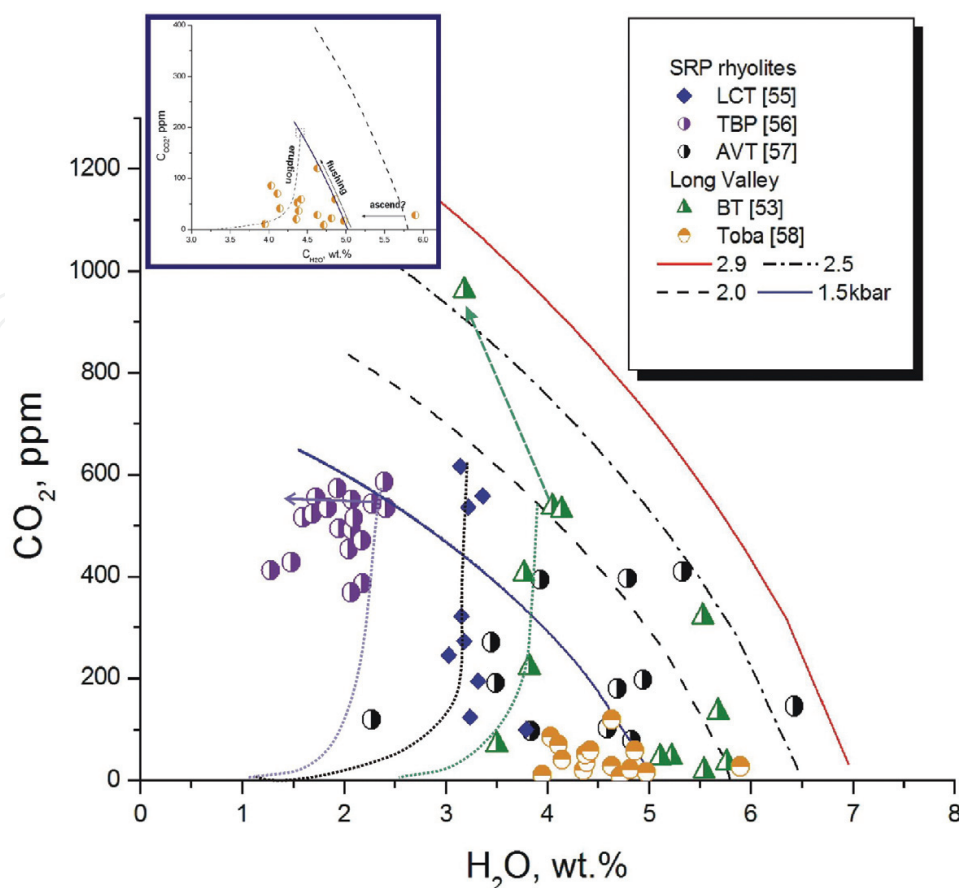


Figure 7. Contents of CO_2 and H_2O in MIs in quartz from silicic volcanic rocks from western USA and Toba. Line show equilibrium mutual solubilities of CO_2 and H_2O at the different pressures. Points located along isobars may correspond to the CO_2 fluxing of the hydrous silicic magma. Pyroclastics from the bishop tuff eruption (Long Valley caldera) experienced more strong interaction with CO_2 than Yellowstone system. In the Toba volcano localized in the subduction system this interaction is weakest.

alignment of some compositional points along the 1.5 kbar isobar. Other points may be located on the close system degassing trajectories, starting with an isobar of 1.5 kbar (see in the inset in Figure 7).

3. The “heavy breath” of the Yellowstone

3.1. Yellowstone breathes because it is alive

The “heavy breath” of the Yellowstone caldera with periodic (time scale of decades) uplift and subsidence with total amplitude of about 0.5 m (see references in Ref. [36]) was attributed to the exsolved water-rich fluid at the solidification [48] or accumulation of two-phase CO_2 - H_2O fluid in the upper 4 km subsurface layer [39]. Recent episode of the fast uplift in 2004–2007 was studied in detail [59] and interpreted as a result of the deep fluid and magma intrusion at the depths of about 7–12 km (horizontal size of expanding object was about 20 km). Pure fluid invasion was also anticipated near Norris Geyser Basin.

Probable connection with CO₂ fluxing as a cause of the slow Yellowstone chamber deformations is confirmed by the approximately fivefold increase in 1978 in the content of “magmatic carbon” in the annual rings of pines from Mud volcano thermal area [60] associated with earthquakes swarm and transition to the subsidence phase of the caldera surface. Certainly to reach shallow geothermal level CO₂ must pass through the mush zone. We will consider mechanical respond of this deep zone on CO₂ pumping that should have inertia larger than shallow reservoirs that can be in pace with several decades period of deformations. Similar mechanical disturbances of the surficial geothermal system (depth 2–3 km) caused by invasion of the fluid from intruded magma in Campi Flegri geothermal field (Italy) dissipate in the time scale of several years [61] by spreading excess fluid horizontally in Darcy’s flow.

If deep CO₂ enters the mush zone by a mechanism more efficient than Darcy’s filtration (e.g., with superheated rhyolites or through fracture zones), one can expect mush zone to expand. A rule of thumb estimate of the maximum value of the overpressure which can be reached in Δt years is $\Delta P = K_{ch} \left(\gamma q_{CO_2} S \Delta t / \rho_f V_{ch} \right)$, where γ is correction factor accounting water extraction from the melt (here 1.2), Δt is time interval of pumping (here 50 years), S – caldera area (2000 km², we use active area less than maximum one delineated by circular fault), V_{ch} is magma chamber volume that can be taken for Yellowstone as 3000 km³/0.3, K is bulk modulus of 20 GPa, ρ_f is density of fluid at $P = 2$ kbar and $T = 850^\circ\text{C}$ it is 600–500 kg/m³. At $\Delta t = 50$ yrs this order of magnitude estimate is 18 MPa. The coefficient γ characterizes incoming fluid volume increase due to water extraction from the melt. As discussed above, its value depends on the actual volume of the melt that reacted with fluid. If interaction is efficient enough and the mass ratio of fluid/reacted melt is small this coefficient can be up to 4, otherwise it is in the range 1.1–1.2.

3.2. Viscoelastic model of flushing

It follows that at the used parameters values overpressure and associated deviatoric stresses can be high enough for mechanical failure and fluid release followed by the caldera subsidence and beginning of the new cycle. Considering the periodic expansion and contraction of the magma chamber in the time scale of decades, it is important to take into account the viscoelastic properties of heated surrounding rocks. As a first approximation of this problem, the formulation in the form of a viscoelastic spherical shell (approach used in Ref. [62]) can be analyzed. With an increase of the relative shell thickness (the heated and damaged rocks aureole), this approximation at small times approaches the conventional equation for a spherical inclusion in an infinite elastic medium. Because of the spherical symmetry, this problem reduces to 1-D and has a simple analytical solution (see in the textbooks). The geometry of the shell is determined by the outer R_1 and the inner R_0 radii. On the external surface, pressure (radial stress) p_1 and p_0 is applied on the inner surface. It is easy to show that the stress distribution does not depend on the material parameters (elastic, viscous or viscoelastic) and depends only on the shell geometry:

$$s_2 = \frac{-P_1 R_1^3 (R^3 + R_0^3) + P_0 R_0^3 (R_1^3 + 2R^3)}{2R^3 (R_1^3 - R_0^3)}, \quad s_{rr} = -\frac{p_1 R_1^3 (R^3 - R_0^3) + P_0 R_0^3 (R_1^3 - R^3)}{R^3 (R_1^3 - R_0^3)} \quad (15)$$

here s_2 are angle components of stress tensor in spherical coordinates $s_{\phi\phi}=s_{\tau\tau}$. The radial displacement in an elastic solution is

$$u(R) = -P_1 \left(\frac{R^3}{3} + \frac{3K}{4\mu} R_0^3 \right) R_1^3 + P_0 \left(\frac{R^3}{3} + \frac{K}{4\mu} R_1^3 \right) R_0^3 / K(R_1^3 - R_0^3) R^2 \quad (16)$$

With a confining pressure P_1 equal to the internal P_0 , the displacement field is reduces to a uniform displacement field corresponding to the volume deformation and all components of the stress tensor (σ_{rr} , $\sigma_{\phi\phi}$, $\sigma_{\theta\theta}$) becomes $-P$. In the expression for the tangential stress (s_{rr}), the first term, proportional to the confining pressure, is compressive, while the term proportional to the internal pressure is tensile so that $s_{rr} > 0$ for the thin shell and high overpressure. The solution for the radial distribution of displacements $u(R)$ depends on the material parameters and can be used to calculate the deformation of the magma volume due to internal processes. The general viscoelastic solution is easily derived from elastic one through its Laplace transform, obtained by replacing the shear modulus with the complex shear modulus and the boundary pressures on their Laplace transforms. With an instantaneous onset of the pressures P_1 and P_0 ($dP = P_0 - P_1$), the expression for the transient displacements $u(R_1, t)$ at the outer boundary is a sum of

$$\begin{aligned} u_1(R_1, t) &= dP \frac{\alpha^3 R_1 t}{4(1 - \alpha^3)\eta}, \quad u_2(R_1, t) = dP \frac{R_1(1 + \nu)(2 - \nu)\alpha^3}{3E\nu} \frac{1}{2(1 - \alpha^3)}, \quad u_3(R_1, t) \\ &= -dP \frac{R_1(1 - 2\nu)^2\alpha^3}{3E\nu(1 - \alpha^3)} \exp\left(-\frac{3E\nu}{2(1 + \nu)^2\eta} t\right), \end{aligned} \quad (17)$$

where the first term (Eq. (17)) describes viscous deformations of the shell at the applied overpressure dP and two next terms describe viscoelastic deformations. Sum of u_3 and u_2 corresponds to some expansion due to exponential relaxation. Term u_2 corresponds to the residual displacement at the full relaxation. Elastic displacements $u_2 + u_3(t = 0)$ (equivalent to the surface uplift at $dP > 0$) are proportional to dPR_1/E . By setting $R_1 = 3000$ m, $dP = 30$ MPa and $E = 30$ GPa, we get $u = 0.9$ m initial deformations and 1.2 m residual uplift (at $dP = \text{const} = 30$ MPa). Viscosity $\eta = 10^{19}$ Pas can be considered as typical for the weak lower crust below brittle-ductile transition boundary. This boundary arises above active magma chambers. Viscosities $\eta \leq 10^{18}$ Pas delineate zone the closest thermal aureole. One can distinguish between this zone and magma mush only in a short time scale deformations. By setting effective viscosity of the shell of $\eta = 10^{19}$ Pas we get viscous uplift (expansion) rate of 2 cm/y. At time interval 60 years viscous and elastic deformations equated in magnitude. Relaxation time scale at these parameters will be 44 years.

3.3. Solution for magma pressure variation

For the general dependence of the internal pressure $P_0 + dP(t)$ on time, the expression for radial displacements is formulated with viscoelastic and viscous terms (retaining terms depending on the overpressure $dP(t)$) as:

$$u_{ve}(R_1, t) = \frac{R_1 \alpha^3}{(1 - \alpha^3)} \left(dP(t) \frac{3(1 - \nu)}{2E} - \frac{(1 - 2\nu)^2}{2\eta(1 + \nu)^2} \int_0^t \exp \left(-\frac{3E\nu(t - \xi)}{2(1 + \nu)^2\eta} \right) dP(\xi) d\xi \right) \quad (18a)$$

$$u_{ve}(R_0, t) = \frac{R_0 \alpha^3}{(1 - \alpha^3)} \left(dP(t) \frac{1 + \nu + 2\alpha^3(1 - 2\nu)}{2E\alpha^3} - \frac{(1 - 2\nu)^2}{2\eta(1 + \nu)^2} \int_0^t \exp \left(-\frac{3E\nu(t - \xi)}{2(1 + \nu)^2\eta} \right) dP(\xi) d\xi \right) \quad (18b)$$

$$u_{visc}(R_1, t) = \frac{R_1 \alpha^3}{4\eta(1 - \alpha^3)} \int_0^t dP(\xi) d\xi, \quad u_{visc}(R_0, t) = \frac{R_0}{4\eta(1 - \alpha^3)} \int_0^t dP(\xi) d\xi \quad (18c)$$

Relative volume increment (dV/V) in time t then can be expressed through displacement of the inner surface as $dV/V = 3(u_{ve}(R_0, t) + u_{visc}(R_0, t))/R_0$. Eqs. (18) is transformed into nondimensional form with pressure scale P_{sc} (50 MPa) and time scale t_0 (10 years):

$$\frac{dV}{V} = \frac{3u(R_0)}{R_0} = \frac{3\alpha^3}{(1 - \alpha^3)} \left(c_1 dP(\tau) - c_2 \int_0^\tau e^{c_3(\tau - \xi)} dP(\xi) d\xi + \frac{3c_4}{(1 - \alpha^3)} \int_0^\tau dP(\xi) d\xi \right), \quad (19)$$

where coefficients are:

$$c_1 = \frac{1 + \nu + 2\alpha^3(1 - 2\nu)P_{sc}}{2E\alpha^3}, \quad c_2 = \frac{P_{sc}t_0(1 - 2\nu)^2}{2\eta(1 + \nu)^2}, \quad c_3 = \frac{3E\nu t_0}{2(1 + \nu)^2\eta}, \quad c_4 = \frac{P_{sc}t_0}{4\eta} \quad (20)$$

and can be expressed through two independent physical parameters namely, the Deborah number, $De = Et_0/\eta$, and nondimensional Young modulus, E/P_{sc} . A geometric factor $\alpha = R_0/R_1$ is another independent variable. Poisson's ratio, ν , is in the range 0.2–0.25. We use the physical parameters values as in the examples considered above, that is, Young's modulus $E = 30$ GPa, viscosity $\eta = 10^{19}$ Pas, Poisson's ratio $\nu = 0.25$, $\alpha = 0.5$. To get the simplest solution, we take a polynomial representation of the overpressure in nondimensional time:

$$dP(\tau) = \sum_{i=1}^3 a_i \tau^i \quad (21)$$

Upon substituting $dp(\tau)$ in Eq. (19), it takes a closed form. Then, it becomes easy to find an approximation that produces close to linear in time volume growth at a constant fluid flux $dV/V(\tau)$ in the interval $0 < \tau < 5$ by minimizing $\int_0^\tau \left(dV(dp(\xi))/V_{ch} - \gamma q_f \xi/V_{ch} \right)^2 d\xi$. This is easy to do in Maple, and the order of magnitude of the deviation of mean integral square deviation for the third-order polynomial is 10^{-12} .

Relative magma chamber expansion dV/V at $\tau = 5$ ($t = 50$ years) is the controlling parameter depending on the multiplication coefficient γ and CO₂ flux. Indeed, even the chamber volume V_{ch} is not well constrained. We take V_{ch} in the range $(2-3) \times 10^4$ km³, $\gamma = 1.2-4$, $S = 1000-2000$ km², $q_{CO_2} = 40$ kg/m²/yr. and get at $\tau = 5$

$$dV/V = \frac{\gamma \tau q_{CO_2} S}{V_{ch} \rho_{fl}} = 0.5 \times 10^{-4} - 0.002 \quad (22)$$

With variation of q_{CO_2} in the range 0–100 kg/m²/yr., the relative volume increase is within 0–0.005. We calculate the variation of overpressure with time for different parameters values (see **Figure 8a**). For a fixed set of mechanical parameters and different CO₂ fluxes ($dV/V = (0.3\text{--}3) \times 10^{-3}$), the overpressure is 1–11 MPa. Hydraulic fracturing with accumulated CO₂ starts at an overpressure that exceeds the tensile strength of rocks, which is realistically estimated to be around 6 MPa [63]. It means that in 10–50 years, CO₂ will be released into the upper crust and after uplift the caldera will subside. Indeed, mechanical parameters and relative shell thickness also influence the pressurization rate (see **Figure 8b**), and the given calculations cannot be considered as an irrefutable proof for the discussed mechanism of “heavy breath”. Nevertheless, results of these calculations show that for a certain range of parameters values the mechanism can correctly explain the observations. More field data are required to unambiguously distinguish between accumulation of residual fluid at the solidification [48], and CO₂ infilling. The first mode corresponds to the fast reduction of the magma volume while second to its growth or thermal equilibrium.

4. Acceleration of basaltic magma supply rate due to deglaciation

Large caldera-forming eruptions (LCFE) require a special eruptive mechanism, since normally overpressure in the chamber reaches zero when evacuating a very small fraction of the magma volume [63]. Dunson [63] has suggested that if an eruption occurs while the roof simultaneously subsides along outward dipping faults, then the overpressure can remain on a significant level. Indeed, it has been shown [1] that it is enough to have one outside dipping fault crossing the roof of the magmatic chamber to form a caldera with a trap-door configuration, which can later evolve into a full scale piston-like subsidence of the roof. At the early stage of an eruption, a strong overpressure is required to cause a critical failure of the roof of the magma chamber or the activation of a regional fault. Then, the integrity of the entire structure can be lost and pieces of the roof will literally float in the magma pool below. This will cause an abrupt drop in the magma pressure. When all the volume of the magma (or partial melt) is saturated with respect to the H₂O–CO₂ fluid (which is possible with a CO₂ fluxing), vesiculation will start throughout the magma volume and a reduced magma density will have a feedback effect on the pressure. In this scheme, the factor that creates a sharp increase in the rate of magma generation becomes a trigger capable of initiating LCFE.

4.1. Effect of deglaciation on the melting rate

Deglaciation is a factor, accelerating the rate of magma generation and strongly affecting the stress state of the crust. Glaciation in the Northern Hemisphere was initiated c.a. 3 Myr ago (e.g., [64]). Since then, periodically large ice caps of up to several kilometers thickness were formed and melted in the high latitudes. Ice is similar to low density sedimentary rocks however, at the onset of the interglacial, ice melting rate exceeds normal denudation rates thus creating geologically unprecedentedly high unloading rates. Recently, it was shown that deglaciation locally accelerates erosion rate that enhances net unloading effect [65]. In the areas of magma generation caused by decompressional melting in the ascending mantle flows

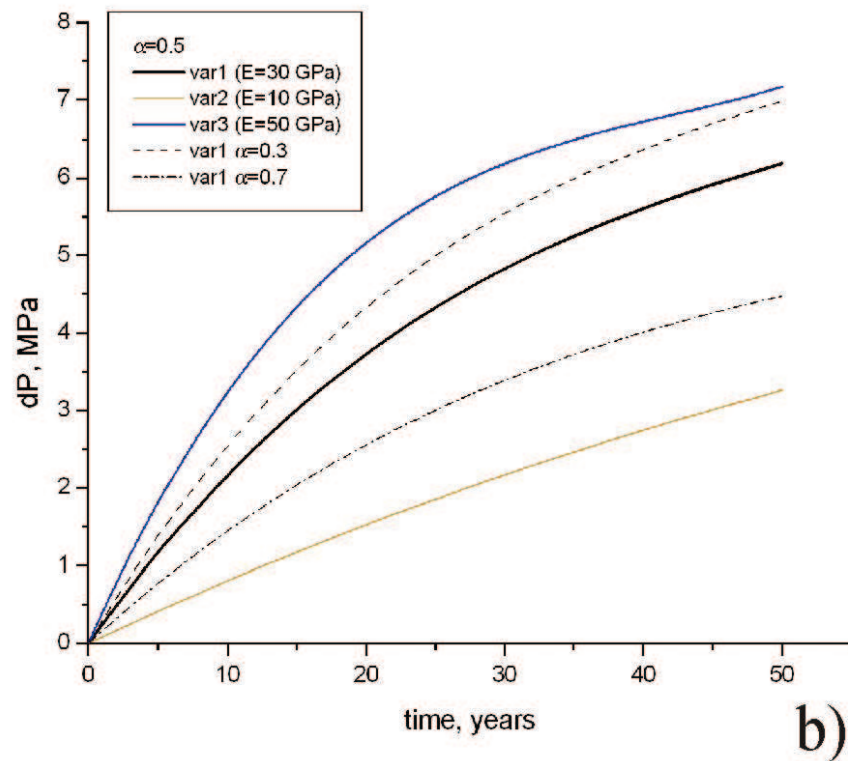
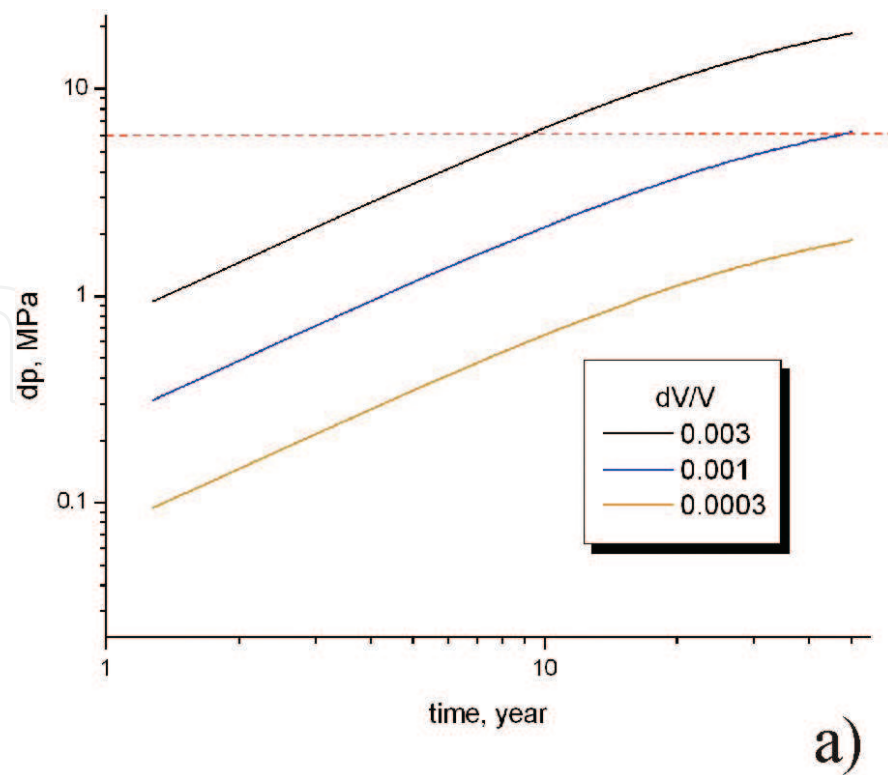


Figure 8. Calculated pressure increase dP at the infilling of magma chamber with CO₂ at $P = 2$ kbar. Simplified model of a spherical viscoelastic shell filled with magma is used (see text). Different inflation rates are applied overpressure is calculated. In 50 years, the magma volume is increased by dV/V (a) dV/V is in the range 0.0003–0.003 (b) $dV/V = 0.001$ and different values of Young's modulus, E , and relative thickness of the shell ($\alpha = (R_1/R_0)^3 - 1$) are used.

(mantle plumes, compensatory flows at the delamination, back arc setting in subduction related mantle upwelling) fast decrease of the glacial load produces significant increase in the melting rate. Decompression melting in the ascending mantle flow occurs at a characteristic pressure decrease rate of $(3-5) \times 10^{-4}$ MPa/yr at a flow velocity of 1.0–1.5 cm/yr. [66]. Thus, it appears that when the glacier melting time is about 1000 years, the associated decompression rate is 0.01 MPa/a. That is, the expected increase in melting rate is 20–30 times. In accordance with this estimate, the volumetric eruption rate in Iceland increased 20 times in the Holocene [67]. The effect of deglaciation on the Yellowstone giant magmatic chamber will be considered next.

4.2. Quaternary glaciations in Yellowstone

Geologic observations in the north-central United States (Iowa, Nebraska, Kansas and Missouri) [68] reveal that some of the tills that marked termination of major glaciation episodes were covered by ashes from the last Yellowstone eruptions (2.1, 1.2 and 0.64 Myr ago). Geographically studied region is far away from Yellowstone but great advance of the ice cap from Canada recorded there probably was in phase with development of glaciation in the highlands in western United States including Yellowstone. Probably multiplicity of caldera-forming eruptions at this location of the mantle plume contrasted with single major eruptions in other calderas of Snake River plain is related to the influence of deglaciation. The last young post-LCT rhyolites eruptions of Central Plateau started in c.a. 260, 176, 124, 79 Kyr [69]. All of them are well correlated with beginning of the global interglacials (see **Figure 9**) after a maximum of 262, 182, 132, 93 Kyr as recorded in the ice core of Vostok site of Antarctica ice shield [70]. Obviously, glacial state of the Yellowstone can differ from the global average reflected in the ice core. Delay of about 14 Kyr for the last event can partially be linked with deviation from the integrated climate variation record. It is also unclear why not all global interglacials have paired rhyolitic eruptions episodes. Christiansen [71] presented arguments that some Central Plateau Member rhyolites were emplaced against glacial ice that may imply that not all eruptions were initiated by deglaciation. These estimates are very preliminary and need extensive field studies to search glacial and postglacial signatures onsite.

4.2.1. Last Pleistocene glaciation in Yellowstone

More is known about the last phase of Pleistocene glaciation in Yellowstone. Based on the fluid inclusions study [48], it has been concluded that Yellowstone plateau was covered with ice sheet with thickness 450–750 m. However, there were no eruptions that can be expected in connection with at least 10-fold increase of the generation rate of basaltic magma. Only hydrothermal explosions producing craters at most 2×3 km in size occurred in early Holocene [72]. First moment after ice load removal is characterized by the largest differential stresses at depths close to the half-width of a glacier. In Holocene, strong earthquakes with magnitude larger than 6 were recognized even in the tectonically passive platforms [73]. Such tectonic events can produce new fractures changing the stress state of the hydrothermal reservoir and provide paths for the deep fluid, thus increasing heat transfer. Hydrothermal explosions are commonly explained by the pressure drop in the water-dominated hydrothermal system with temperature close to boiling [74]. For example, abrupt decrease of the water table due to drainage of glacially dammed lakes [72] will induce deep boiling, reservoir expansion and mechanical disruption with boiling front

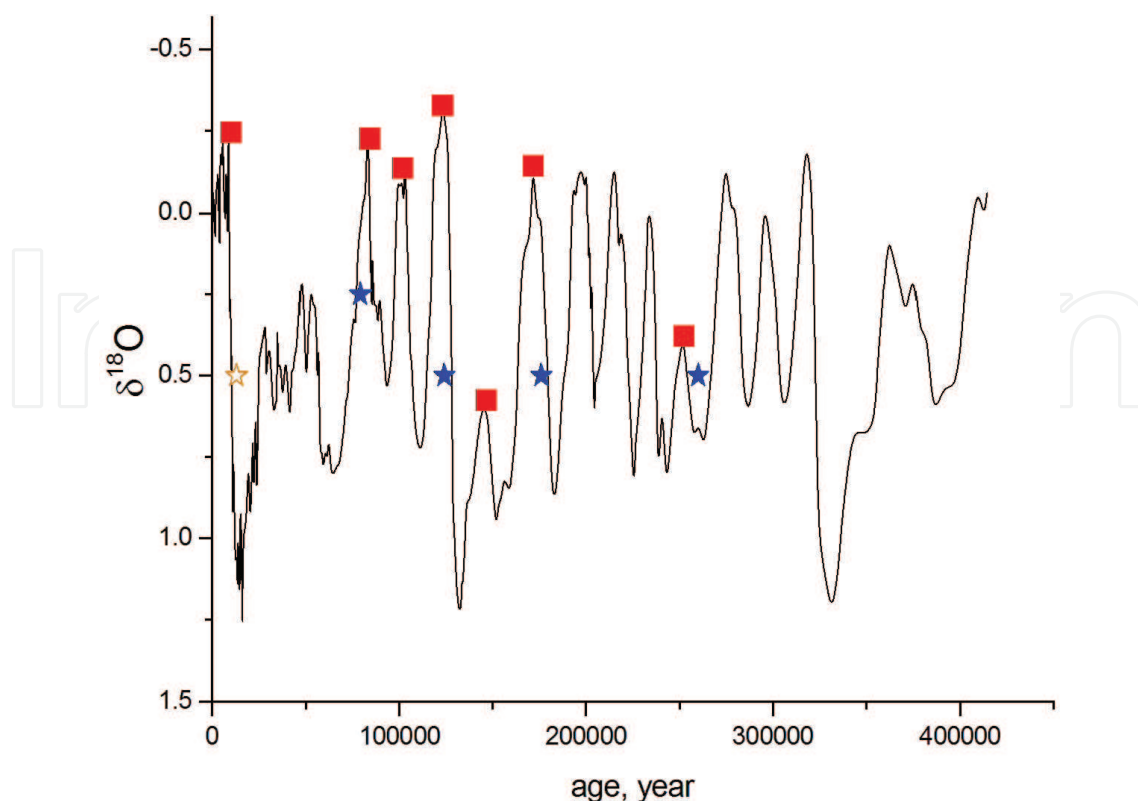


Figure 9. Isotope composition $\delta^{18}\text{O}$ of the oxygen from air in ice the core from the Vostok station, in East Antarctica [70] is plotted as global glaciation index, red squares are at the total ice mass minima (interglacials). Blue stars indicate time of the beginning of the recent volcanic activity in Yellowstone [69]. Semi-filled circle corresponds to the burst of the hydrothermal explosions in Holocene [72].

propagating downwards. In their review Browne, Lawless [74] noted that accumulation of gaseous CO₂ near the surface may enlarge the effect of hydrothermal explosion in classic mode. Similar effects can be produced by the fast invasion of the CO₂-rich hot fluid into water-dominated hydrothermal reservoir. A similar mechanism was proposed by Hedenquist and Henley [75], suggesting that Waiotapu (New Zealand) geothermal system was under stress when injected by magmatic gases that induced strong hydrothermal explosion (crater diameter 5 km, maximum depth of ejected clasts 350 m).

4.2.1.1. Some geochemical evidences for CO₂ fluxing in Yellowstone

Reduced carbonic fluid is a good solvent not only for Pt but for Sr as well as other elements [76]. Thus, one can expect that CO₂ fluid passing through the mush zone would transfer light rare earth elements (LREE), Sr from the basalt magma to rhyolitic magma and extract Ba, Sr and some other elements from it, carrying them to the near-surface hydrothermal system.

Our suggestions are consistent with the geochemical data on hydrothermally altered rhyolites from Mary Bay (Yellowstone Lake) [72]. These altered rocks are enriched in Ba (up to 18 times) and especially Sr (up to 80 times) in comparison with intracaldera rhyolites, as expected if their transfer by deep CO₂ is significant. The Sr concentrations in the altered pyroclastics are much higher than in the extracaldera basalts [23]. More interestingly, these rocks (former rhyolites)

have a high Ni content up to 1600 ppm. A close positive correlation between the content of Ni and Co, Ni and Cr and Ni and V was encountered. These observations suggest that some organometallic-compounds of these siderophilic elements were present in the fluid. A detailed hydrological study [77] revealed a clear outflow of Ba and Sr (extracted from the rhyolite mush) into the Yellowstone Lake with geothermal fluids.

4.2.1.2. *Current state of Yellowstone magma chamber*

In the view of the high uncertainties in the estimates of the heat and CO₂ fluxes and basaltic magma supply rate, results of the glacial probes which reflect the integrated state of the magma chamber under Yellowstone are of exceptional importance. Mechanical test of the last glaciation demonstrated that currently, the magmatic system of Yellowstone is far from the impending catastrophic eruption. Impulse of the basaltic magmatism inevitably accompanying last deglaciation was dumped in the massive mush zone with low partial melt content. It is noteworthy that the time interval of rhyolite eruptions in the post LCT series [69] monotonously decreased with time, which can be interpreted as manifestation of the decay in the basaltic magma generation under the caldera. Moreover “periodic deglaciation forced evacuation” of magma from Yellowstone magma chamber might prevent it from a major eruption.

5. Conclusions

Since extra-large magma chambers require extra-large volumes of the added magma in a short time to erupt, eruptions have an extremely low frequency. Some large magma chambers, such as Yellowstone, are located in areas not only of exceptional magma generation rate, but also of very high CO₂ inflow which can significantly affect the mechanical stability of the chamber.

Deep CO₂ flux in the measured range alone cannot prevent the cooling and solidification of the Yellowstone mush zone. In combination with the thermal energy of basaltic magma, a temporary increase in the flow of CO₂ can switch the thermal state from the thermal equilibrium to the mush melting.

Based on the analysis of melt inclusions, it is possible to estimate the current integrated mass flux of CO₂ to be at least 0.1–0.3 of the total melt mass in the mush zone of Yellowstone. The geochemical consequence of the long CO₂ fluxing can be averaging the magma composition and reducing of the content of Sr and Ba extracted into the geothermal system.

Simple modeling of deformations of viscoelastic rocks of the thermal aureole of Yellowstone magma chamber herein, shows that deep CO₂ fluxing in a reasonable range of flow rate and mechanical parameters can lead to slow uplift and subsidence cycles with a full amplitude of about 1 m and a period of several decades.

The main triggering mechanisms of the eruption are spikes in the rate of generation of basaltic magma caused by internal processes in the mantle plume underlying Yellowstone. In the Quaternary, repeated glaciation in the Northern Hemisphere became another factor that periodically increased magma generation rate 10–30 times in early interglacials. The last Pleistocene

glaciation in Yellowstone caused only hydrothermal and not magmatic eruptions in the Holocene. Since the previous interglacials were marked by voluminous post-LCT rhyolite eruptions, the result of Holocene glacial probing may imply that the current state of the Yellowstone magma system is in a stable or decaying thermal regime.

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